

Support Vector Machine Practice

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Hyperparameter

- A model hyperparameter is a configuration that is external to the model and whose value cannot be estimated from data.
 - They are often used in processes to help estimate model parameters.
 - They are often specified by the practitioner.
 - They can often be set using heuristics.
 - They are often tuned for a given predictive modeling problem.

Revisit Soft Margin

- Maximize the number of samples 1) while the minimizing the number of samples of 2) & 3)
- We can revise the hard margin example into

$$\text{Minimize } \frac{\|w\|^2}{2}$$

$$g(\mathbf{x}_i) = 1 - y_i(\mathbf{w}^\top \mathbf{x}_i + b) \leq 0$$



$$\text{Minimize } \frac{\|w\|^2}{2} + C \sum_{i=1}^N \xi_i$$

$$\text{subject to } 0 \leq \xi_i$$

$$g(\mathbf{x}_i) = 1 - \xi_i - y_i(\mathbf{w}^\top \mathbf{x}_i + b) \leq 0$$

- We can repeat the procedure by the Lagrange aux. function of

$$L = \frac{\|w\|^2}{2} + C \sum_{i=1}^N \xi_i - \sum_{i=1}^N \mu_i \{y_i(\mathbf{w}^\top \mathbf{x}_i + b) - 1 + \xi_i\}$$

Revisit Famous Kernel Function

- Radial basis function or **RBF** kernel or Gaussian kernel

$$\kappa(\mathbf{x}, \mathbf{x}') = \exp\left(-\frac{\|\mathbf{x} - \mathbf{x}'\|^2}{2\sigma^2}\right)$$

- Polynomial kernel

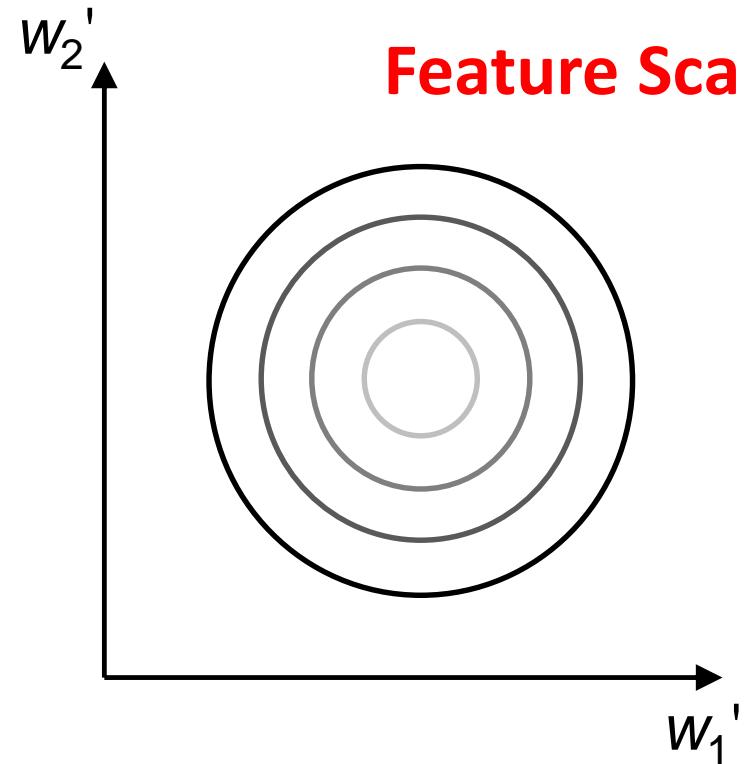
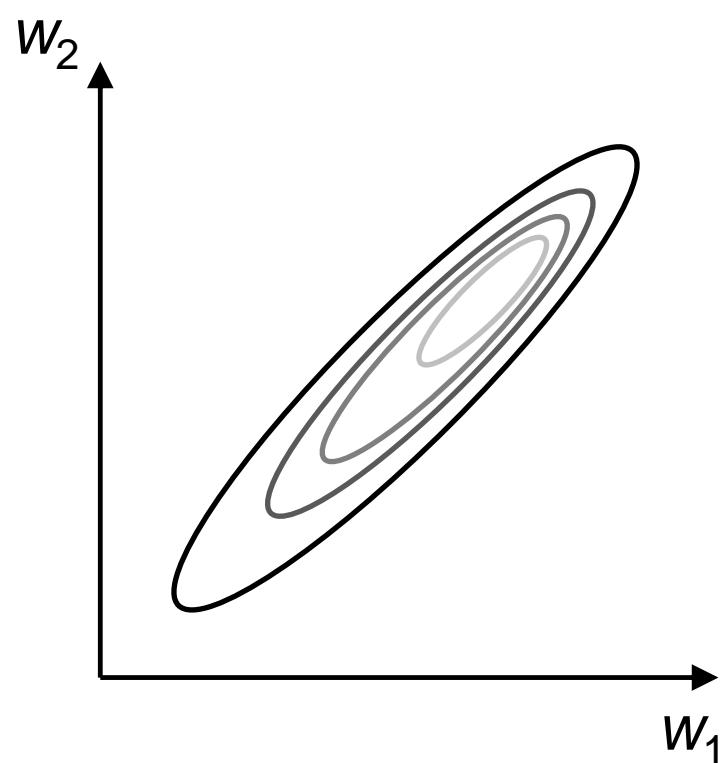
$$\kappa(\mathbf{x}, \mathbf{x}') = (\gamma \mathbf{x}^T \mathbf{x}' + r)^M$$

- Sigmoid kernel or tanh kernel

$$\kappa(\mathbf{x}, \mathbf{x}') = \tanh(\gamma \mathbf{x}^T \mathbf{x}' + r)$$

Gradient Descent in Different Data Scale

$$w_{\text{next}} = w_{\text{present}} - \eta \nabla NLL(w)$$



Feature Scaling in Python

- Min-max scaling (Normalization)

- Make the data limited to [0,1]
 - Vulnerable to outlier noise

- Standardization

- Does not limited to certain range
 - Robust to outlier noise

- Usage:

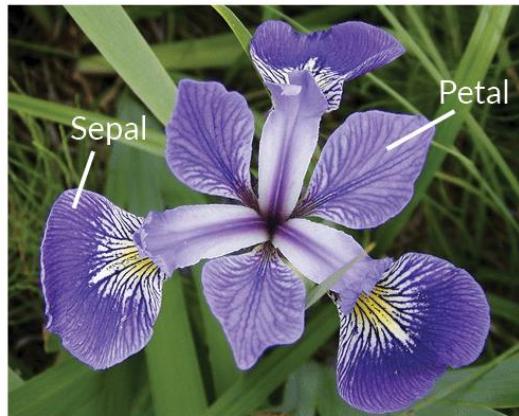
```
from sklearn.preprocessing import StandardScaler
```

```
X = np.array([
    [2, -3],
    [4, 1],
    [0, -2],
    [10, 3]])
```

```
scaler = StandardScaler()
scaler.fit(X)
X_std = scaler.transform(X)
```

Load iris Data for SVM Practice

- Iris data?



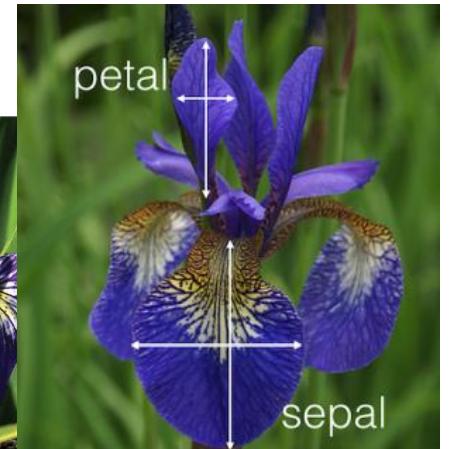
Iris Versicolor



Iris Setosa



Iris Virginica



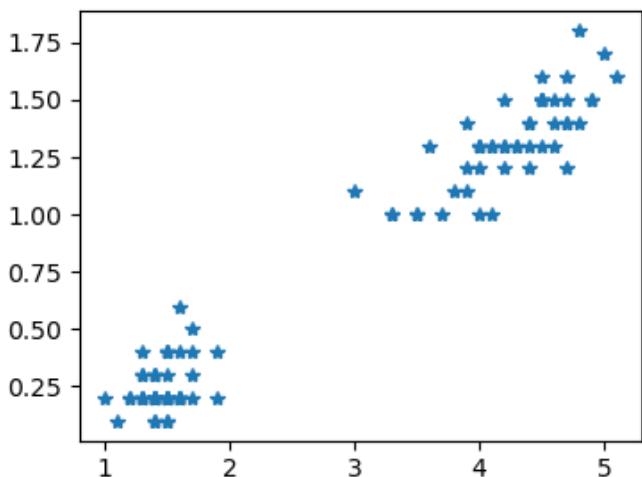
```
import matplotlib.pyplot as plt
from sklearn import datasets
iris = datasets.load_iris()

X = iris["data"][:, (0, 1)] #petal length & width
Y = iris["target"][:100]

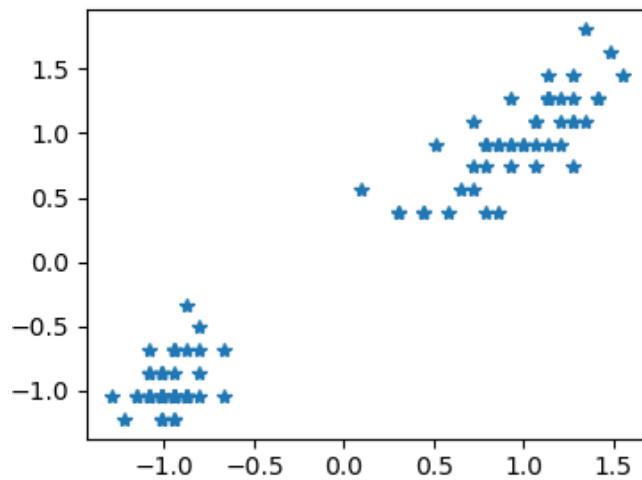
f1, ax1 = plt.subplots()
ax1.plot(X[:, 0], X[:, 1], '*')
```

Classified iris Datasets w/ Standardization

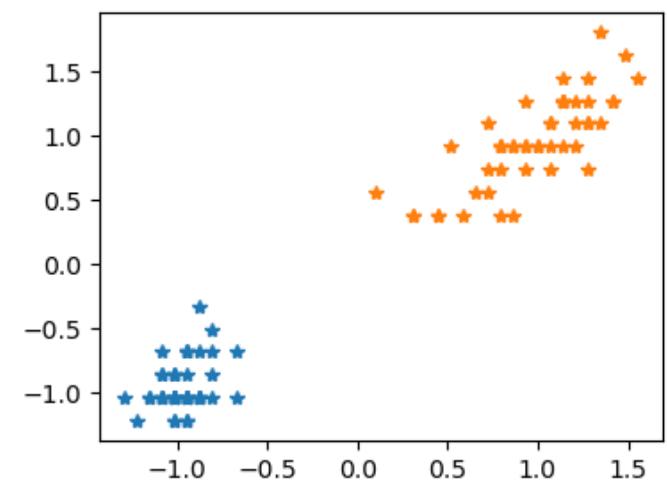
iris petal width vs. petal length



Standardization



Classified



```
import numpy as np
import pandas as pd
import matplotlib.pyplot as plt

from sklearn import datasets
from sklearn.preprocessing import StandardScaler

plt.close("all")

iris = datasets.load_iris()
X = iris["data"][:, 2:4] #petal length & width
Y = iris["target"][:100]
f1, ax1 = plt.subplots()
ax1.plot(X[:,0], X[:,1], '*')

scaler = StandardScaler()
scaler.fit(X)
X_std = scaler.transform(X)
f2, ax2 = plt.subplots()
ax2.plot(X_std[:,0], X_std[:,1], '*')

df_clf = pd.DataFrame(np.c_[X_std, Y])
df_clf.columns = ["petalLength", "petalWidth", "target"]
df_clf_group = df_clf.groupby("target")
f3, ax3 = plt.subplots()
for target, group in df_clf_group:
    ax3.plot(group.petalLength, group.petalWidth, '*', label = "target")
```

SVM by sklearn

- Although we can simply write a code for SVM by realizing:

$$L(\mu) = \sum_{i=1}^N \mu_i - \frac{\sum_{i=1}^N \sum_{j=1}^N \mu_i \mu_j y_i y_j \mathbf{x}_i^\top \mathbf{x}_j}{2}$$

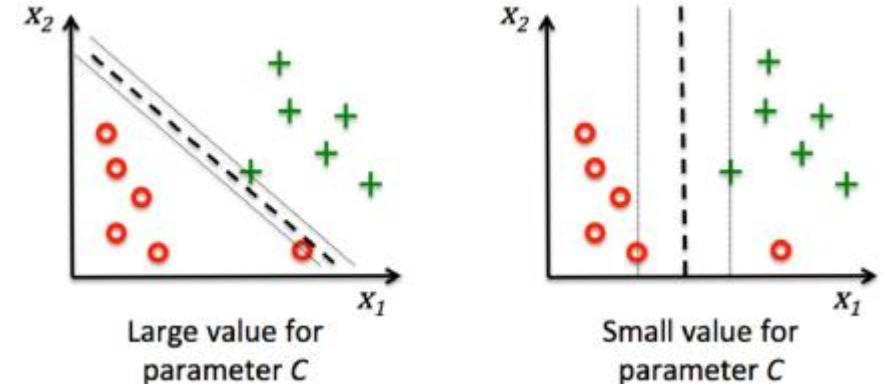
Then, w & b from optimal μ

- We can utilize the pre established code:
- [sklearn.svm.SVC — scikit-learn 0.24.2 documentation](#)

Usage of SVM

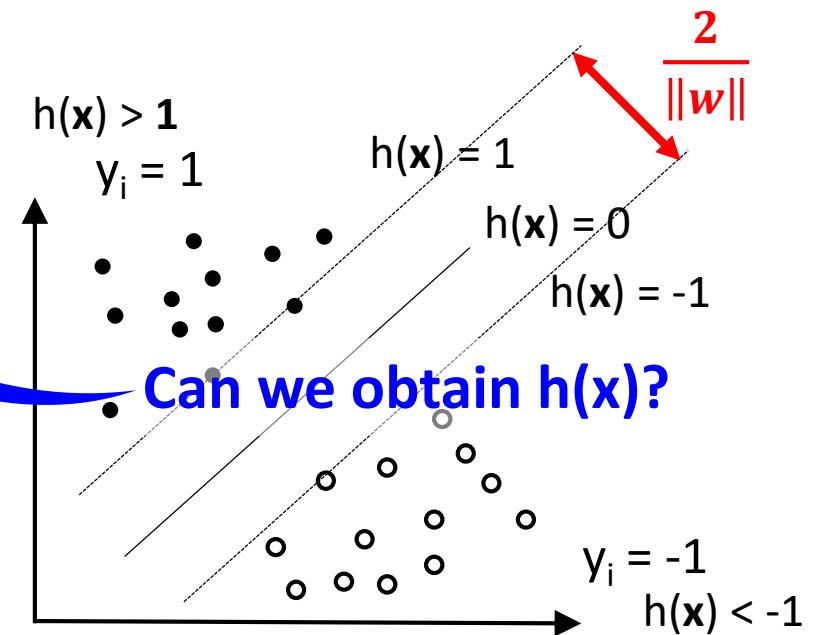
```
svm_clf = SVC(C=0.01, kernel="linear")
svm_clf.fit(X_std, Y)
```

```
fit(X, y, sample_weight=None)
Fit the SVM model according to the given training data.
```



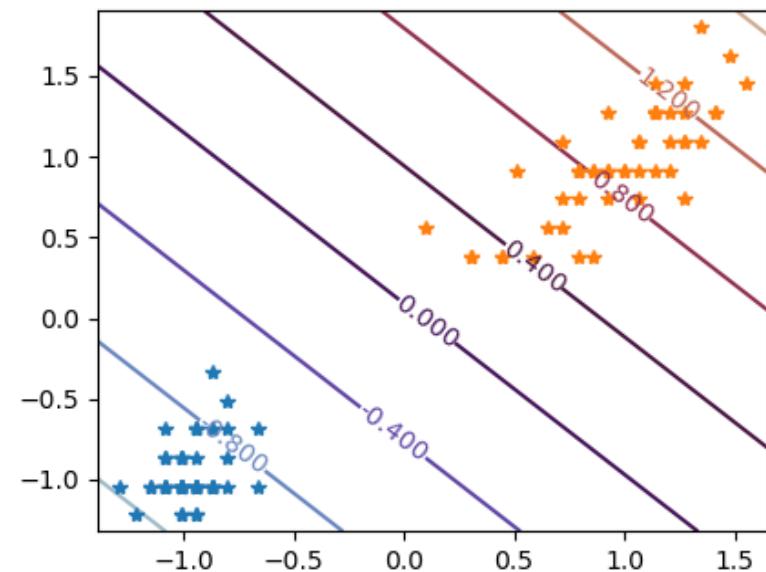
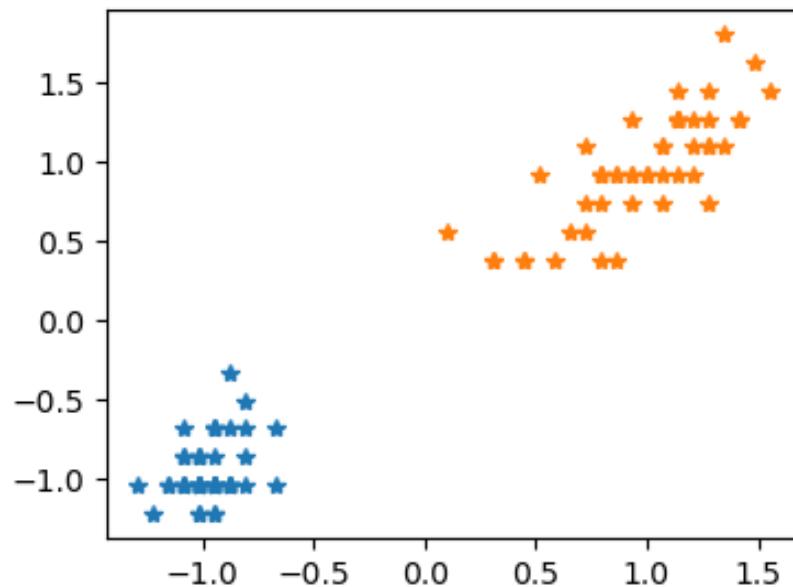
- Now, check `svm_clf` in Variable Explorer.
- What is matter for us?

`svm_clf.decision_function(X)`



Visualization of SVM Training Results

- We have `svm_clf.decision_function(X)`
- How can we visualize this?
- We can utilize contour with meshgrid

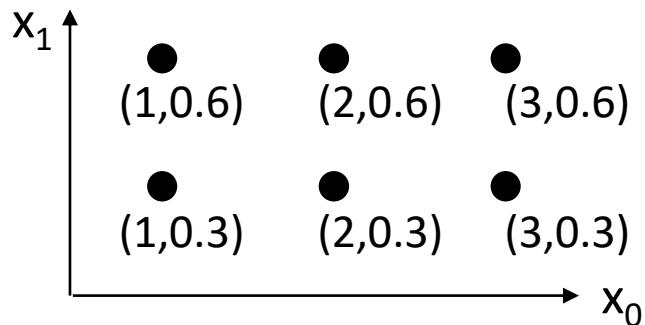


Contour

- `ax3.contour(x0Plt, x1Plt, h)` How should `x0Plt` or `x1Plt` look like?

`[h(1, 0.3), h(2, 0.3), h(3, 0.3)
h(1, 0.6), h(2, 0.6), h(3, 0.6)]`

`[1, 2, 3] [0.3, 0.3, 0.3
1, 2, 3] 0.6, 0.6, 0.6]`



How can we make if we have
 $x0 = [1, 2, 3]$ and $x1 = [0.3, 0.6]$?

`[x0Plt, x1Plt] = np.meshgrid(x0, x1)`

$x0$
`[1 2 3]`
 $x1$ `[0.3 1, 0.3 2, 0.3 3, 0.3
 0.6] 1, 0.6 2, 0.6 3, 0.6`



$x0Plt$ $x1Plt$
`[1, 2, 3] [0.3, 0.3, 0.3
1, 2, 3] 0.6, 0.6, 0.6]`

How Can We Achieve $h(X)$?

x0Plt	x1Plt	
[1, 2, 3]	[0.3, 0.3, 0.3]	→
1, 2, 3]	0.6, 0.6, 0.6]	

[$h(1, 0.3), h(2, 0.3), h(3, 0.3)$
 $h(1, 0.6), h(2, 0.6), h(3, 0.6)$]

We need [[1, 0.3]
[2, 0.3]
[3, 0.3]
[1, 0.6]
[2, 0.6]
[3, 0.6]]

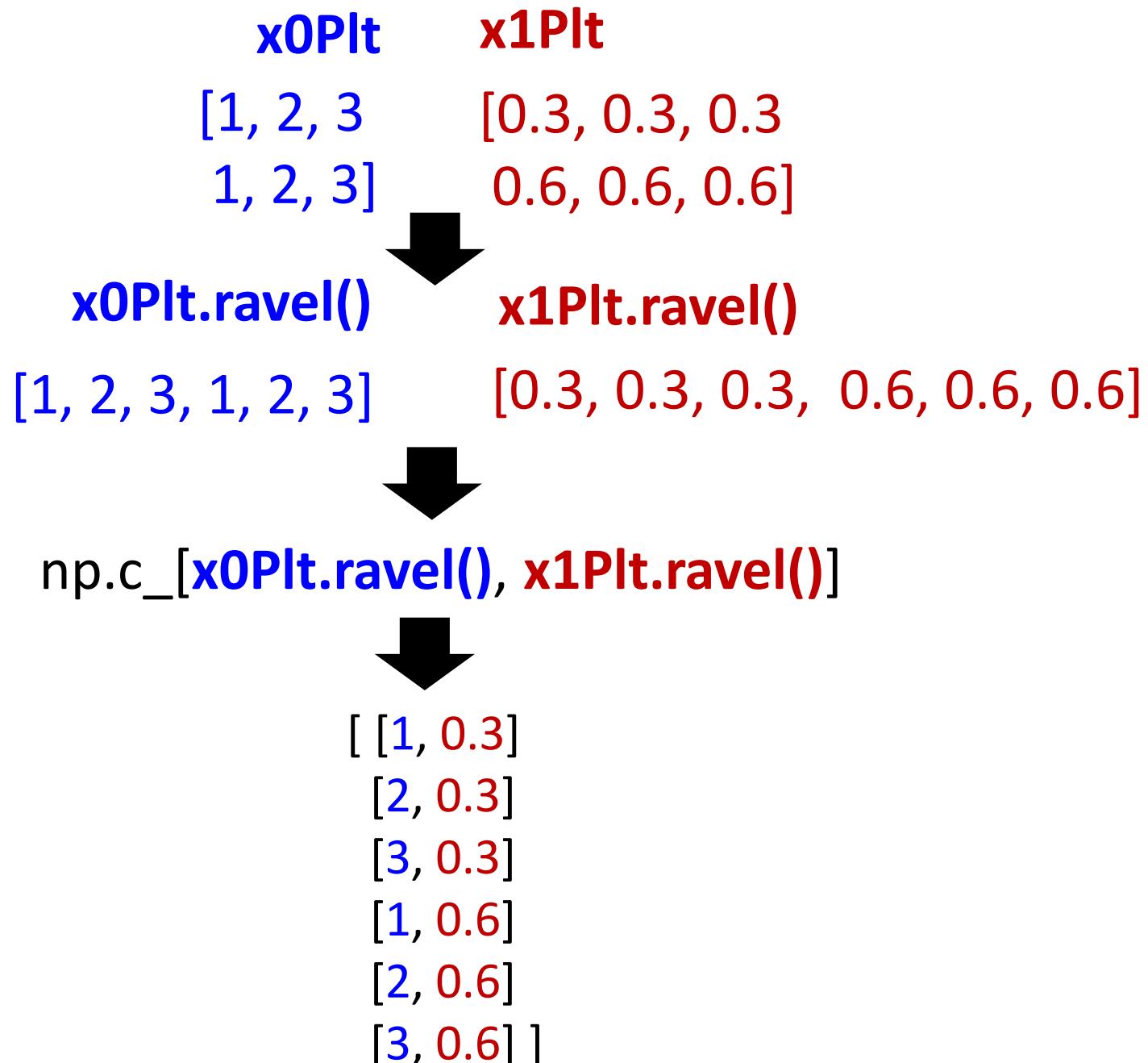
Reshape vs. Ravel

```
a = np.array([0, 1, 2, 3, 4, 5, 6, 7, 8, 9, 10, 11])
```

a.reshape(3,4)

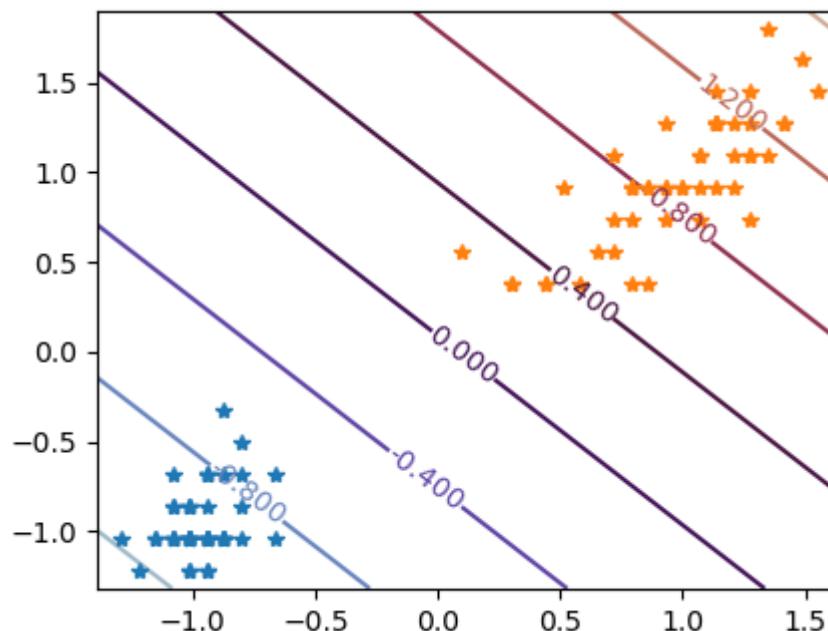
a.ravel()

```
a = np.array([[0, 1, 2, 3], [4, 5, 6, 7], [8, 9, 10, 11]])
```



Code for Visualizing Contour

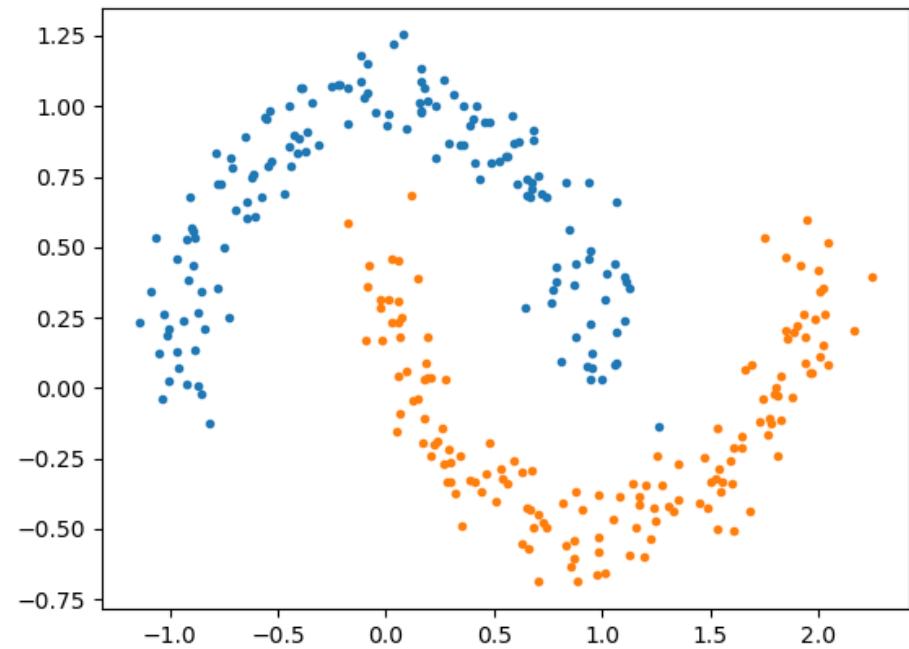
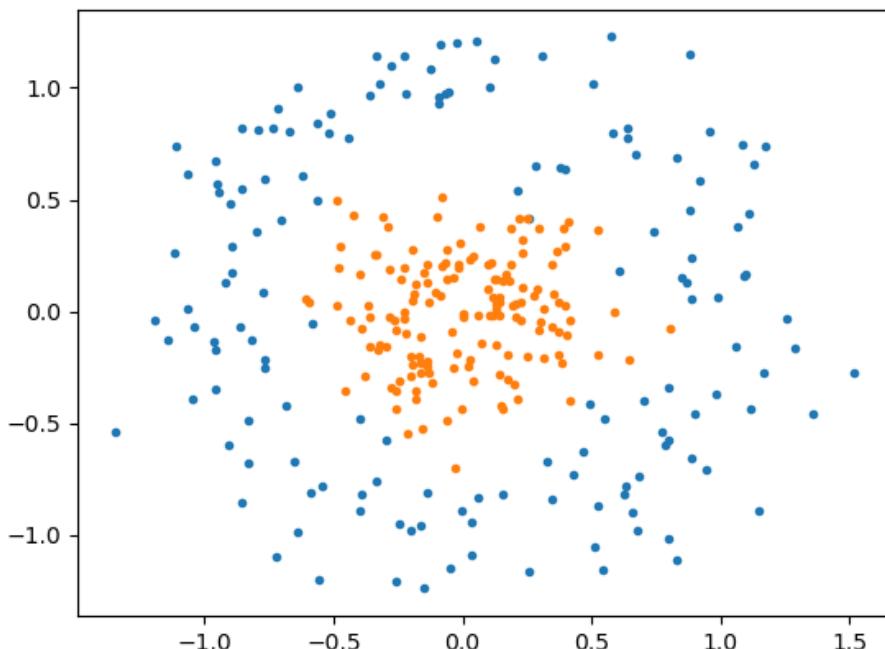
```
[x0Min, x0Max] = [min(X_std[:,0])-0.1, max(X_std[:,0])+0.1]
[x1Min, x1Max] = [min(X_std[:,1])-0.1, max(X_std[:,1])+0.1]
delta = 0.01
[x0Plt, x1Plt] = np.meshgrid(np.arange(x0Min, x0Max, delta),
np.arange(x1Min, x1Max, delta))
h = svm_clf.decision_function(np.c_[x0Plt.ravel(), x1Plt.ravel()])
h = h.reshape(x0Plt.shape)
CS = ax3.contour(x0Plt, x1Plt, h, cmap=plt.cm.twilight)
ax3.clabel(CS)
```



You also can use
`svm_clf.predict([[x1, x2]])`

Moon-Data or Circle-Data

```
from sklearn import datasets  
plt.close("all")  
# [X,Y] = datasets.make_circles(n_samples = 300, shuffle = True ,  
noise = 0.2, random_state = 15, factor = 0.3)  
[X,Y] = datasets.make_moons(n_samples = 300, shuffle = True,  
noise = 0.1, random_state = 15)
```



Do it By Yourself!

- Apply “train_test_split” to test the SVM performance
- Apply kernel trick to data for **moon data** and **circle data** by changing
 - Kernel function → rbf and poly
 - For rbf, change gamma
 - For poly, change degree/gamma/coef0
 - Repeat above various factor and noise of moon/circle data

$$\kappa(\mathbf{x}, \mathbf{x}') = \exp\left(-\frac{\|\mathbf{x} - \mathbf{x}'\|^2}{2\sigma^2}\right)$$

$$\kappa(\mathbf{x}, \mathbf{x}') = (\gamma \mathbf{x}^T \mathbf{x}' + r)^M$$